Search for star-forming region in the extreme outer Galaxy with WISE data

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Introduction I : EOG Extreme outer galaxy (EOG) : *R*_G ≥ ~ 18 kpc

~ 18 kpc

NASA/JPL-Caltech

- Environment

 Small or no perturbation from the spiral arm
 Low gas density



3. Low metallicity



Introduction I : EOG
Why are we interested in the EOG ?
1. Unexplored region
→ Determine the size and shape of our Galaxy

 Excellent laboratory for studying star-forming processes in low-density and low-metallicity environment

3. Similar characteristics with dwarf galaxy and the early phase of the formation of our Galaxy

Furguson et al. 1998, Kobayashi et al. 2008

We may be able to **directly observe "galaxy formation processes" in unprecedented detail** at much closer distance than High-z galaxies

Our objective

- Examine star-formation properties in the EOG
 - Star formation rate (SFR)
 - Star formation efficiency (SFE)
 - SFR/SFE appear to decrease significantly under low-density / low-metallicity environments



Sun et al. 2014

Our objective

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Should confirm if such trend is present in our Galaxy in molecular cloud scale (pc scale)

Search for young star-forming regions in the EOG (w/ embedded clusters)

Past studies

- ► Mostly based on IRAS color e.g., Brand and Wouterloot 1995、Snell et al. 2002
- Only few star-forming regions are confirmed in the EOG
 - (Sensitivity-limited ?)

Kobayashi & Tokunaga 2000 Brand & Wouterloot 2007

• WISE

- Great increase in sensitivity (100 times more sensitive than IRAS)
- Very high potential in searching for distant star-forming regions



Fabinsky, 2006

• Our target : CO molecular clouds

There are many surveys of CO molecular clouds in recent years





Search for young star-forming regions in the EOG (w/ embedded clusters)

CO survey (II and III quadrant)



New surveys with even higher sensitivity are ongoing

- Nobeyama 45m Galactic plane survey (Japan)
- · Milky Way Imaging Scroll Painting Project (China)

NASA/JPL-Caltech

1) Solar neighbourhood

- $D \leq 1 \text{ kpc}$ (clusters in Lada & Lada 2003)



2) Near outer Galaxy

- **D** = 1 ~ 3 kpc (clusters in Lada & Lada 2003)



3) Outer Galaxy

- **D** = 4 ~ 8 kpc (HII regions in Fitch et al. 1984)



4) Far outer Galaxy

 D = 7 ~ 10 kpc (HII regions in Fitch et al. 1984 clusters in Snell et al. 2002)



Representative cases

- Low-mass cluster (M_{cluster} ~ 10² M_☉)
- Resolved or unresolved ?
 - Typical distance between cluster members :
 Δ= 2-3' at d=150 pc (case for Taurus)
 - WISE resolution : 6-12"



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Star-forming regions in the EOG

- Only three are known
 - Digel Cloud 1 : $R_G = 22 \text{ kpc} (D = 16 \text{ kpc})$
 - Digel Cloud 2 : $R_G = 19 \text{ kpc} (D = 12 \text{ kpc})$
 - WB89-789 : $R_G = 20 \text{ kpc} (D = 12 \text{ kpc})$

Izumi et al. 2014 Yasui et al. 2006. 2008, Kobayashi et al. 2008 Brand & Wouterloot 2007



Izumi et al. 2014

Yasui et al. 2008

Brand & Wouterloot 2007

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Yasui et al. 2008

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WISE View of Star-Forming Region in the EOG Star-forming regions in the EOG

- Digel Cloud 1, 2

Izumi et al. 2014, Yasui et al. 2006, 2008



WISE View of Star-Forming Region in the EOG Star-forming regions in the EOG

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WISE View of Star-Forming Region in the EOG **Star-forming regions in the EOG** - Digel Cloud 1, 2 Izumi et al. 2014, Yasui et al. 2006, 2008 **Our observations** High resolution ¹²CO mapping Nobeyama 45 m radio telescope ↑ Beam size : 17" Deep NIR (J, H, Ks) imaging Dame et al. Subaru 8.2 m telescope (CfA survey ↑ Limiting mag : 22.0(J), 21.5(H), 20.5(Ks) mag, 5σ $R_{\rm G} = 22$ kpc May et al. 1997 Sun (most distanticloud/ cluster in our Galaxy Nakagawa et al. 2005 **Digel Cloud 2** Digel et al. 1994 (Digel Clouds)

Digel Cloud 1

Brand et al. 1994

B89-789

NASA/JPL-Caltech

Digel Cloud 1

- Embedded clusters are detected in the two CO peaks (Cloud 1a Cloud 1b)



- Cloud 1a (N_{cluster} = 18, M_{cluster} ~ 27 M $_{\odot}$)



Cluster is detected as a few point-like red sources

- Cloud 1b (N_{cluster} = 48, M_{cluster} ~ 28 M_☉)



Cluster is detected as a few point-like red sources

Digel Cloud 2

- Embedded clusters are detected in the two CO peaks (Cloud 2N Cloud 2S)



- Cloud 2N (N_{cluster} = 72, M_{cluster} ~ 43 M_☉)



Cluster is detected as an extended sources

- Cloud 2S (N_{cluster} = 66, M_{cluster} ~ 40 M_☉)



Cluster is detected as a few point-like red sources

Case for Digel Cloud 1, 2

- Star-forming regions are clearly detected with WISE in all four bands

→ WISE can detect star-forming regions even at R_G = 22 kpc

However, cluster members are unresolved
 → We must rely only on integrated flux to pick up star-forming regions with WISE data.

Search for star-forming region in the EOG with WISE data How can we find distant star-forming regions with WISE ?

 Construct detection criteria of distant star-forming regions using WISE colors of confirmed star-forming regions with NIR imaging Search for star-forming region in the EOG with WISE data AIIWISE sources in the Cloud 1 field $(1^{\circ} \times 1^{\circ})$ -SN > 5 (w1, w2, w3), SN > 3 (w4)0 Confirmed star-forming regions Sources within molecular clouds (within 3σ contour) ന 200 Field stars Nobeyama 45m) 00 0 5' 00 0 0

Search for star-forming region in the EOG with WISE data AllWISE sources in the Cloud 2 field (1° × 1°) - SN > 5 (w1, w2, w3), SN > 3 (w4) Confirmed star-forming regions

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- Sources within molecular clouds (within 3σ contour)
- Field stars

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1%

Color - color diagram (w1-w2 vs w2-w3)



Color - color diagram (w1-w2 vs w2-w3)



Color - magnitude diagram (w1 vs w1-w2)



Color - magnitude diagram (w1 vs w1-w2)



Color - magnitude diagram (w1 vs w1-w2)



Comparison with Koenig's diagram

- Color-color diagram for individual YSOs



Koenig et al. 2014 Fig.10

Comparison with Koenig's diagram

- Color-color diagram for individual YSOs



Koenig et al. 2014 Fig.10





WISE Properties of the star-forming region in the EOG

- Star-forming regions appear to have the same color as individual YSOs.
 - But we checked only for two molecular clouds.



- Next, we are going to confirm this results with star-forming regions in the FOG ($D = 7 \sim 10$ kpc, $R_G = 13.5 \sim 18$ kpc)
 - There are 11 embedded clusters confirmed by NIR imaging in the FOG Snell et al. 2002



NIR Confirmed star-forming region
 – ¹²CO (FCRAO survey data : beam size ~ 45"

Color-color diagram of the clusters in the FOG



Color-color diagram of the clusters in the FOG





Location of Star-forming region



Location of Star-forming region



Location of Star-forming region



Star-forming region area

Detection criteria of distant (unresolved) star-forming region

1. Located within CO molecular clouds on the sky

2. Located at Star-forming region area in the color-color diagram

- w1 w2 ≥ ~ 0.5
- w2 w3 ≥ ~ 2.0
- w2 w3 ≦ ~ 6.0

- Detection criteria of distant (unresolved) star-forming region
 - Contamination Problems
 - Ex) Planetary nebula, Background AGNs ...

Koenig et al. 2014

Wright et al. 2010

- Detection criteria of distant (unresolved) star-forming region
 - Contamination Problems
 - Ex) Planetary nebula, Background AGNs ...
 - The probability would be quite low in view of the small area of the distant molecular clouds.

→ Utilize high-resolution CO data to define the cloud region more clearly

Search for star-forming region in the EOG with WISE data **Detection criteria of distant (unresolved)** star-forming region - Contamination Problems FCRAO 14m (resolution ~ 45") Nobeyama 45m (resolution ~ 17") small Cloud 1 e the 5

Search for star-forming region in the EOG with WISE data **New candidates in the EOG** (3.4: blue, 12:green, 22:red) Digel Cloud5 N Digel Cloud5 W Digel Cloud3 8 Digel Cloud6 E Digel Cloud6 S Digel Cloud5 S 0 Digel Cloud8 N Digel Cloud7 Digel Cloud8 S **O New candidate** ¹²CO (Nobeyama 45m data : beam size ~ 17") We found 17 new candidates for star-forming region

Summary and Future Works

Summary

 WISE has great sensitivity to detect distant starforming regions even in the EOG (R_G ≥ 18 kpc)
 All star-forming regions could be detected up to the edge of the Galaxy

- We defined tentative detection criteria using WISE color-color diagram, which may be able to <u>effectively pick up distant star-forming regions.</u>

Future works

 Search for star-forming regions in <u>all molecular</u> <u>clouds in the EOG</u> from the recent CO surveys.

 Study star-formation properties (SFR, SFE) in the EOG with WISE data